

CCSU  
DEPARTMENT OF MATHEMATICAL SCIENCES

COLLOQUIUM

Friday, November 21, 2014

2:00 – 3:00 PM

Maria Sanford, Room 101

LAPLACE'S NEBULAR HYPOTHESIS

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(Joint work with Vasily Yanchilin)

**Abstract:** Recall that, according to the nebular hypothesis, stars form in massive and dense clouds of [molecular hydrogen](#) — [giant molecular clouds](#) (GMC). These clouds are gravitationally unstable, and matter coalesces within them to smaller denser clumps, which then rotate, collapse, and form stars. The protoplanetary disk is an [accretion disk](#) that feeds the central star. Initially very hot, the disk later cools in what is known as the [T tauri star](#) stage; here, formation of small [dust](#) grains made of [rocks](#) and ice is possible. The grains eventually may coagulate into kilometer-sized [planetesimals](#). If the disk is massive enough, the runaway accretions begin, resulting in the rapid—100,000 to 300,000 years—formation of Moon- to Mars-sized [planetary embryos](#). Near the star, the planetary embryos go through a stage of violent mergers, producing a few [terrestrial planets](#). The last stage takes approximately 100 million to a billion years.

In this presentation, we criticize this Laplace's hypothesis by showing that protoplanetary disc may not collapse to a planet. This statement follows from the following basic result we are going to prove.

**For satellites with circular orbit, the orbit of joint energy is always below the orbit of joint impulse momentum.**

The presentation should be accessible for everybody who took Calculus I.

**For further information:**

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